

faken during a season of minimum variation in surface albedo patterns, when the solar planetocentric longitude (L<sub>S</sub>) with respect to Mars was 65° to 75° (with the exception of small areas near long 360° compiled with an L<sub>S</sub> of 89°). All pictures were taken through clear or red filters when the solar zenith angle was small, during Viking spacecraft orbital revolutions 583 through 687 (March and April of 1978) and 1323 through 1378 (February and March of 1980).

ADOPTED FIGURE The figure of Mars used for computing the map projection is an oblate spheroid (flattening of 1/192) with an equatorial radius of 3,393.4 km and a polar radius of 3,375.7 km.

**PROJECTION** The Mercator projection is used for the two maps in this set. Longitudes increase to the west in accordance with astronomical convention for Mars. Latitudes are areographic. CONTROL

Planimetric control for the 1:5,000,000-scale maps used to compile the bases for the two maps was derived from photogrammetric triangulation using Mariner 9 pictures (Davies, 1973). This control net was upgraded through the use of Viking data (Davies and others, 1978). At least 85 percent of the image control points lie within 0.5 mm of the positions established in 1978.

MAPPING TECHNIQUE The mapping bases used for this series were assembled from 1:5,000,000-scale shaded relief maps (Batson and others, 1979), reduced and digitally transformed to the Mercator projection where necessary. During shaded relief portrayal, features on these bases were used to position details taken from Viking Orbiter pictures. Shaded relief is shown as if illuminated from the west; methods and interpretation techniques used were described by Inge (1972) and Inge and Bridges (1976). Details of the albedo markings were taken from medium-resolution Viking Orbiter images and drawn on a photographic copy of the original shaded relief map. The overall albedo balance was maintained by reference to a lowresolution solar albedo map made with Viking Infrared Thermal Mapper (IRTM) data (Kieffer and others, 1981; Pleskot and Miner, 1981).

Shaded relief analysis and representation were made by Barbara J. Hall. Analysis and portrayal of albedo markings were made by Susan L. Davis. No attempt was made to duplicate the color of the Martian surface although the color used may approximate it.

1:15,000,000 series; center of map lat 0°, long 270°; shaded relief and markings (A), nomenclature (N).

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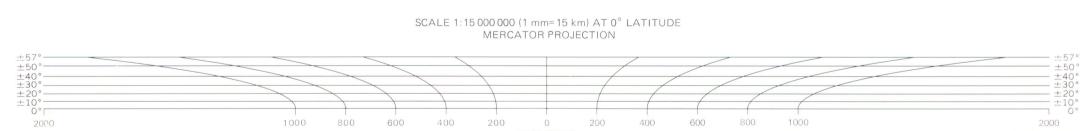
Nomenclature, in 16th General Assembly,

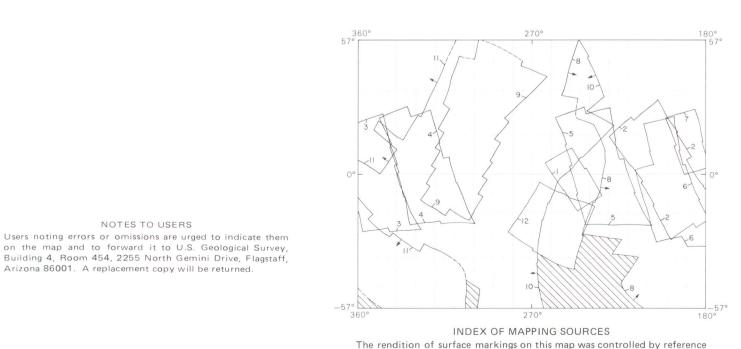
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321-325, 331-336, 355-362. — 1980, Working Group for Planetary System Nomenclature, in 17th General Assembly, Montreal, 1979, Proceedings: International Astronomical Union Transactions, v. 17B, p. 293-297.

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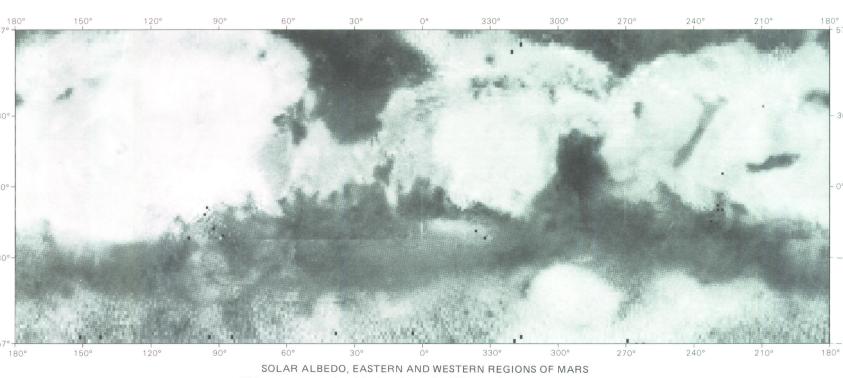




The rendition of surface markings on this map was controlled by reference to the pictures outlined above. Picture numbers consist of an orbital revolution number, a spacecraft designation number, and a frame number. For example, 605A86 was the 86th picture taken by Voyager 1 (A) during revolution number 605. Pictures taken by Voyager 1 after revolution 999 are sometimes designated without the leading digit of the revolution number and with an "S" for the spacecraft number. Thus, 1378A86 is the same picture as 378S86. Dashed lines indicate position of limb. Useful coverage is not available in cross-hatched areas. Copies of various enhancements of these pictures are available from National Space Science Data Center, Code 601, Goddard Space Flight Center, Greenbelt, MD

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This map was made with data from the Viking Infrared Thermal Mapper (IRTM). Broadband brightness (0.3 to 3.0 microns) measured by the IRTM was reduced to Lambert albedo for four Martian seasons that were relatively free of atmospheric dust and local brightening. These data were averaged into resolution elements of 1° latitude by 1° longitude. Absolute albedos range from 0.10 to 0.33 (Pleskot and Miner, 1981).



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